



Completely random reactor model

Egor Stadnichuk^{1,2,3}, Alexander Nozik^{1,2}, Maxim Dolgonosov^{3,4}, Mikhail Zelenyi^{1,2,3}

MIPT¹, INR RAS², SRI RAS³, HSE⁴

Contents

- Model concept
- Simplified model assumptions
- Equation
- Solution and reactor explosion criteria
- Monte Carlo simulation results
- Experimentally observable model properties

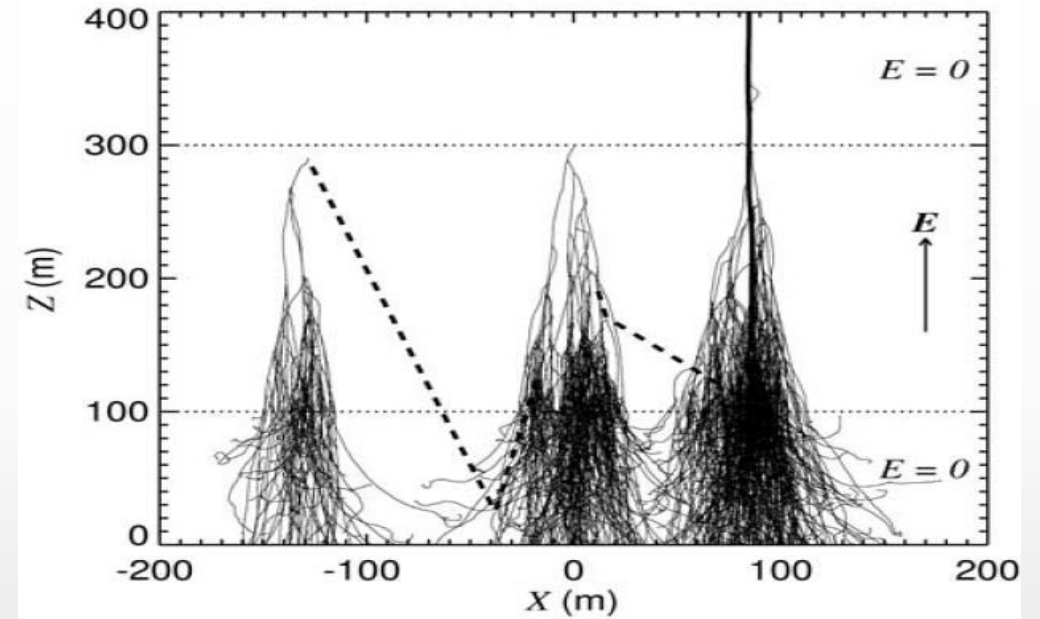


Dwyer RFDM

RREA reproduce themselves by positron and gamma feedback mechanisms.

Positrons are created by RREA bremsstrahlung and propagate in direction opposite to RREA propagation. Consequently, positrons create secondary avalanches at the beginning of the cell.

Does not work in observable conditions.



R. Dwyer J. A fundamental limit on electric fields in air // Geophysical Research Letters. — Vol. 30, no. 20. — URL: <https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2003GL017781>.



RL-TGE model concept

- Thundercloud electric field structure is considered to be complex:

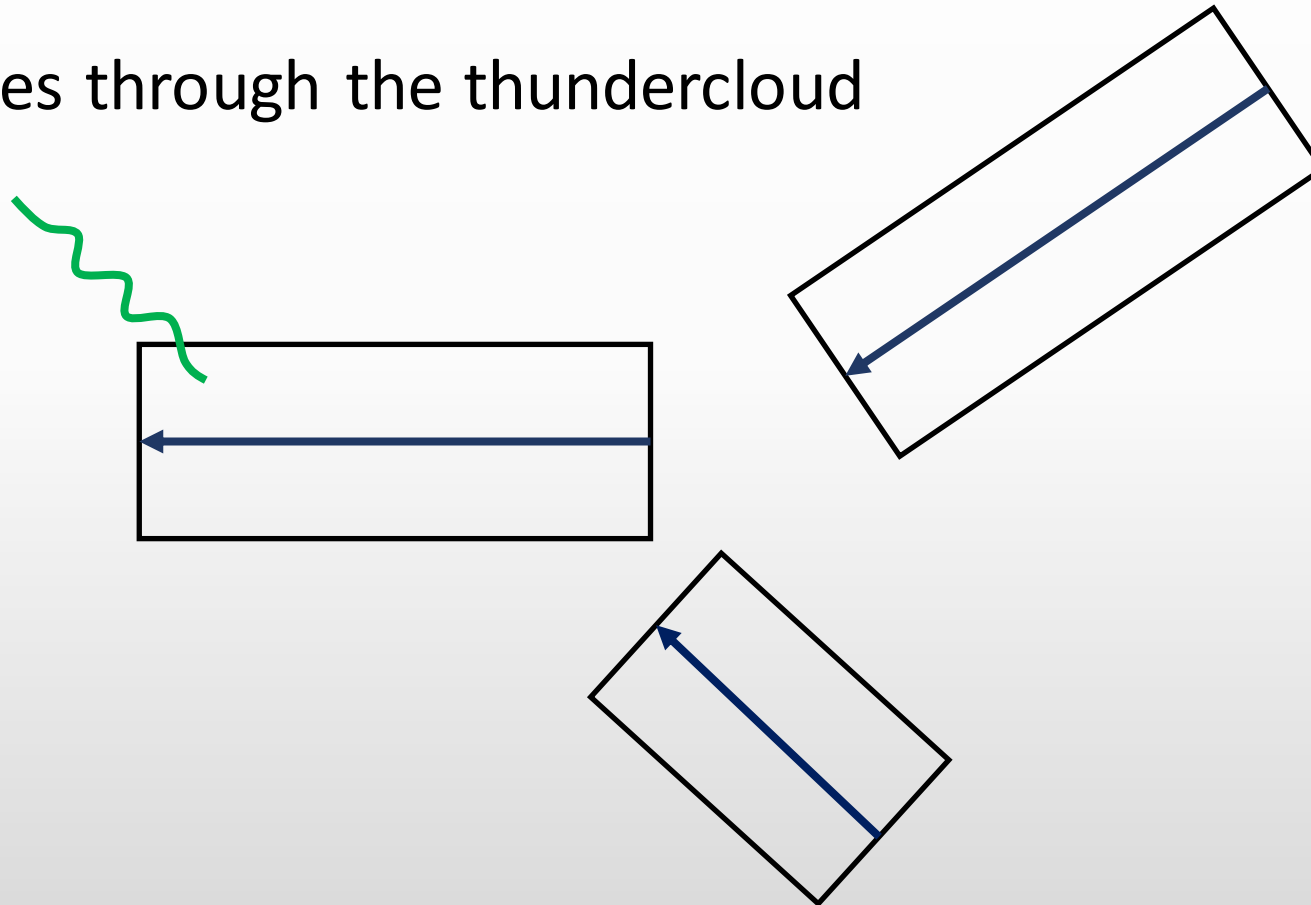
There are a lot of cells with critical electric field and different electric field direction

- Gamma dynamics is similar to neutron dynamics in nuclear reactor:
 - ✓ A gamma propagates through the thundercloud
 - ✓ It gives birth to a runaway electron within a cell
 - ✓ Runaway electron produces RREA
 - ✓ RREA radiates new gammas
 - ✓ These gammas propagate through the thundercloud



RL-TGE model concept

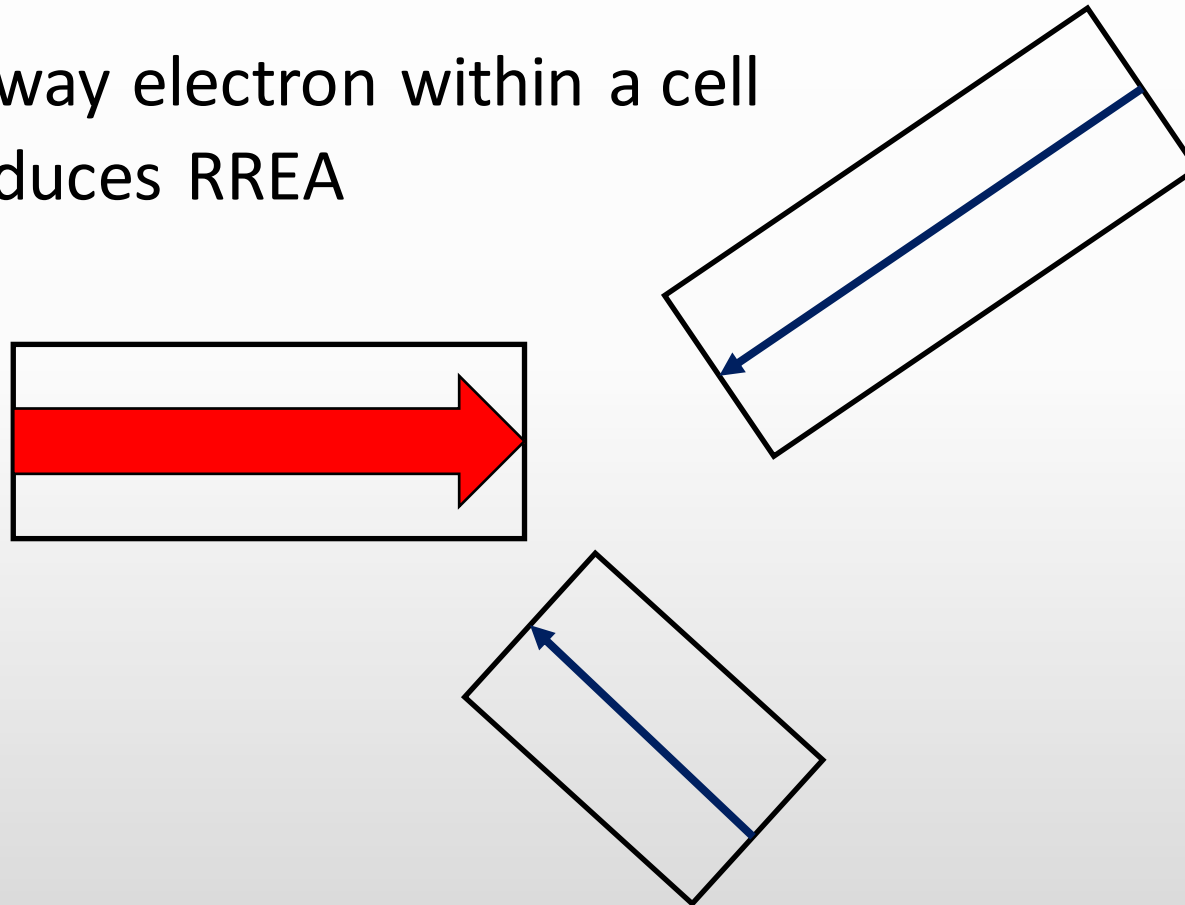
A gamma propagates through the thundercloud





RL-TGE model concept

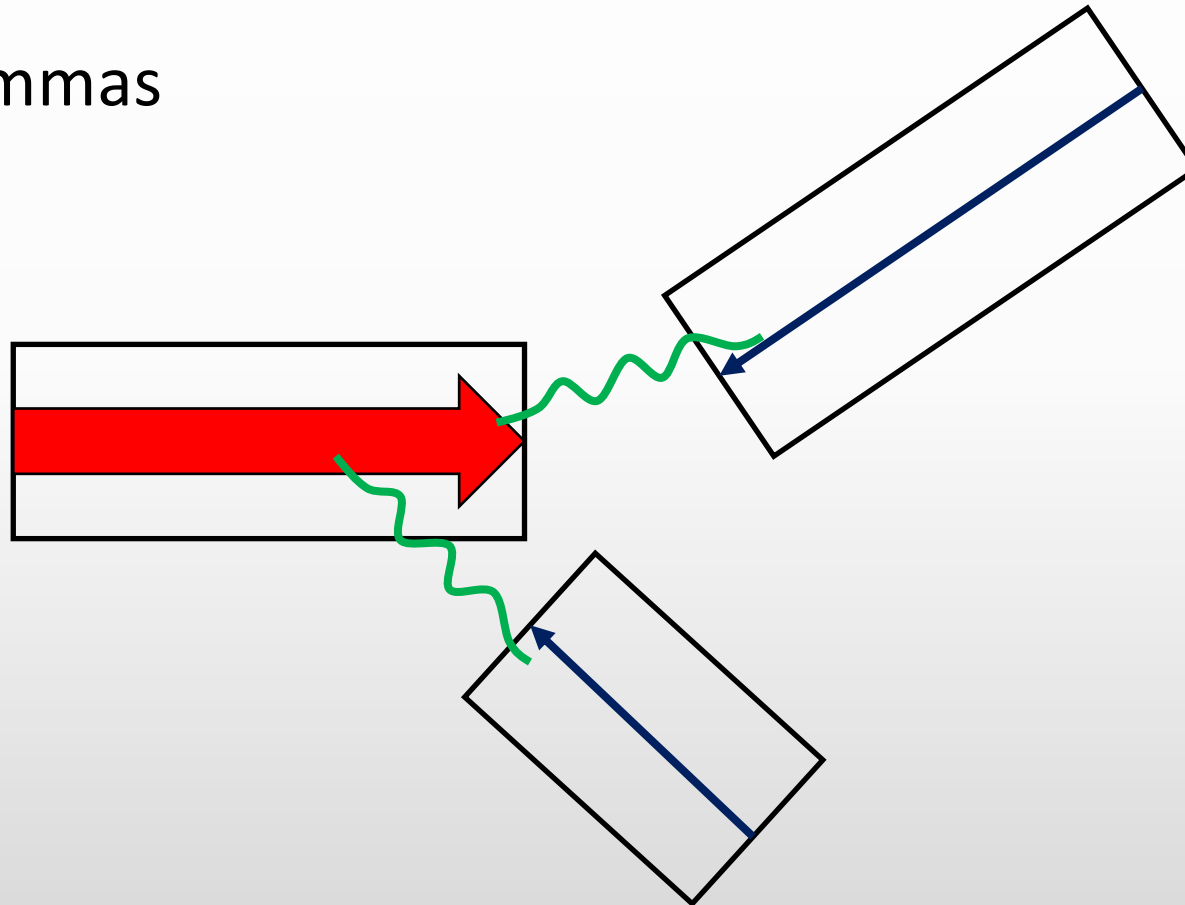
It gives birth to a runaway electron within a cell
Runaway electron produces RREA





RL-TGE model concept

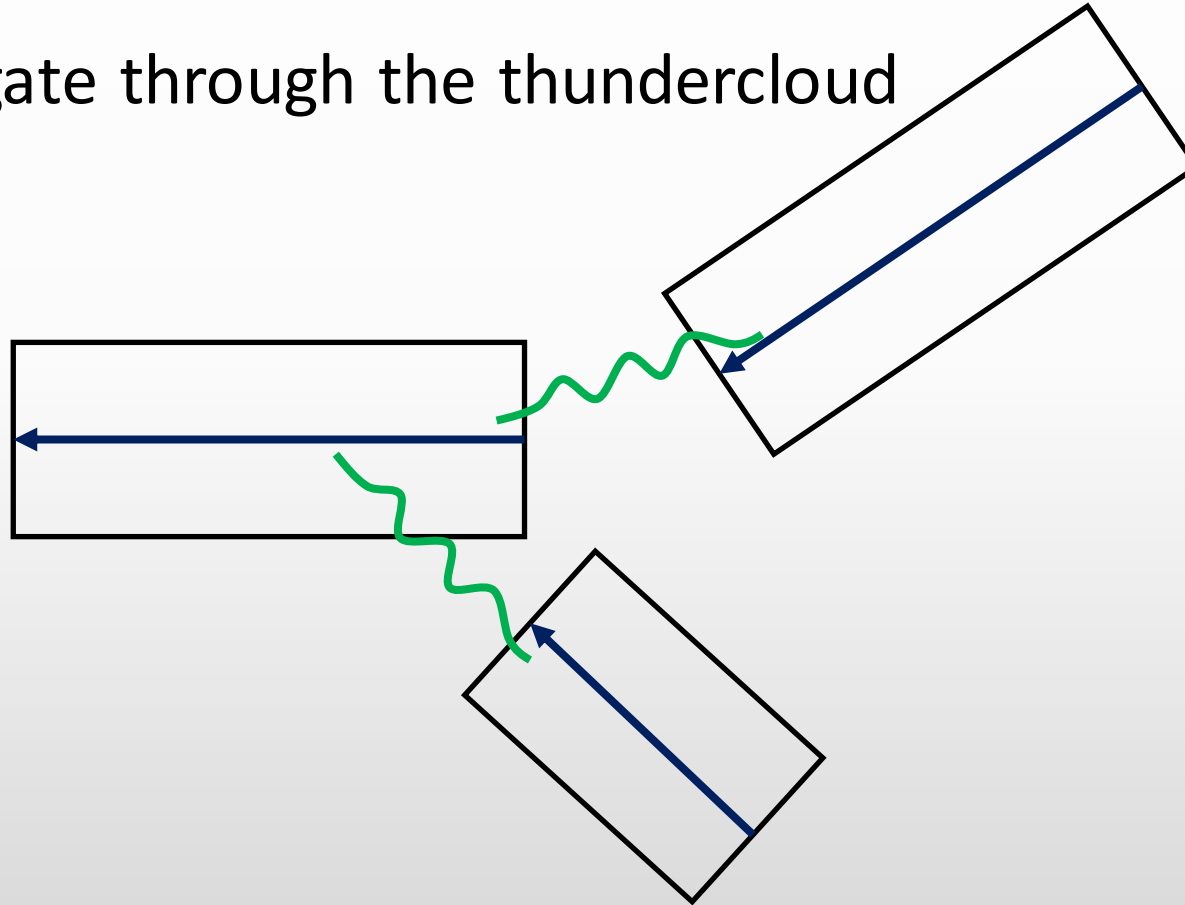
RREA radiates new gammas





RL-TGE model concept

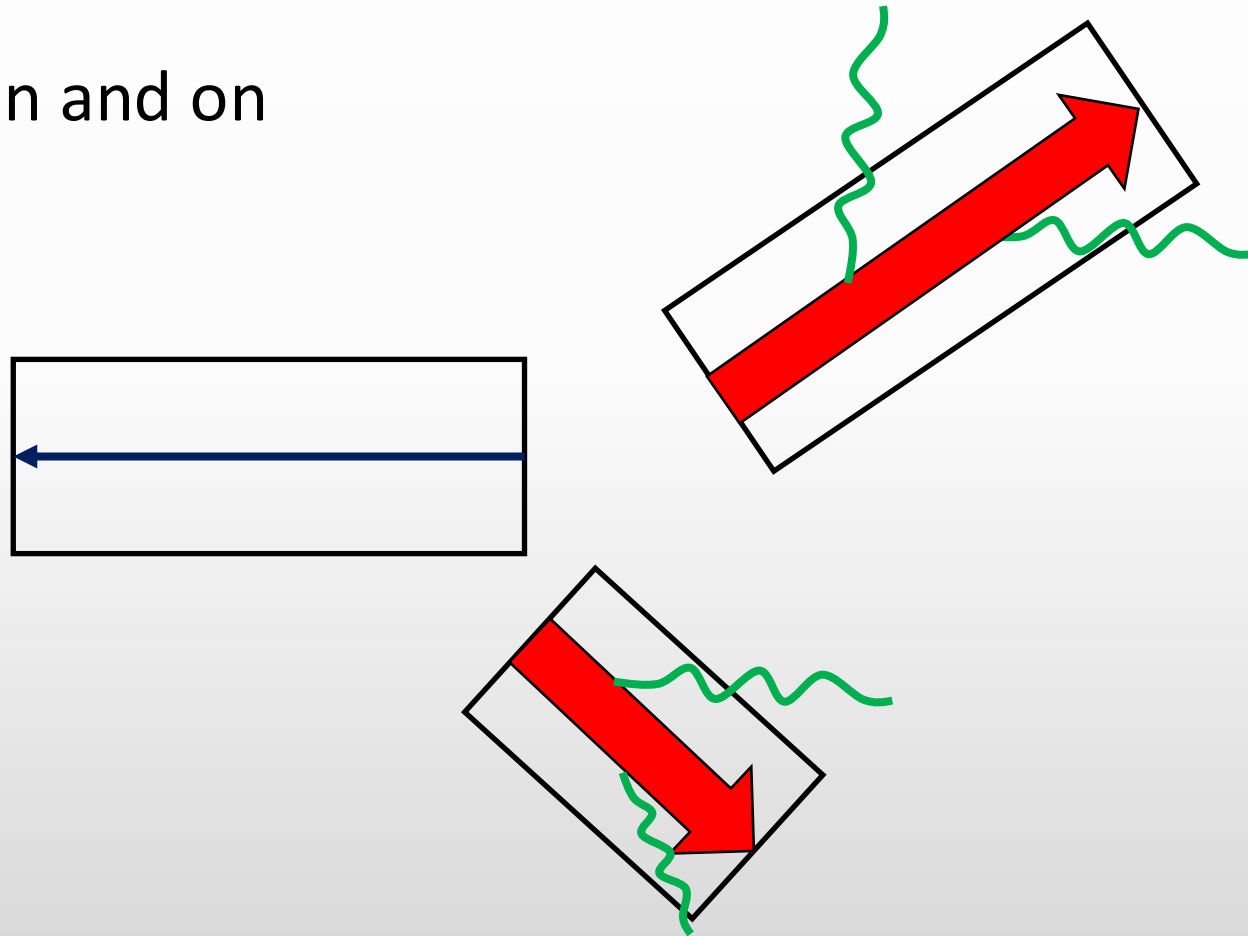
These gammas propagate through the thundercloud





RL-TGE model concept

And everything goes on and on





Reactor model simplified simulation

1 / 51

cell-length

cloud-size

field-magnitude

free-path

gain

particle-limit

output

seed

save-plot

seed-photons

dynamic-plot

Show Cube

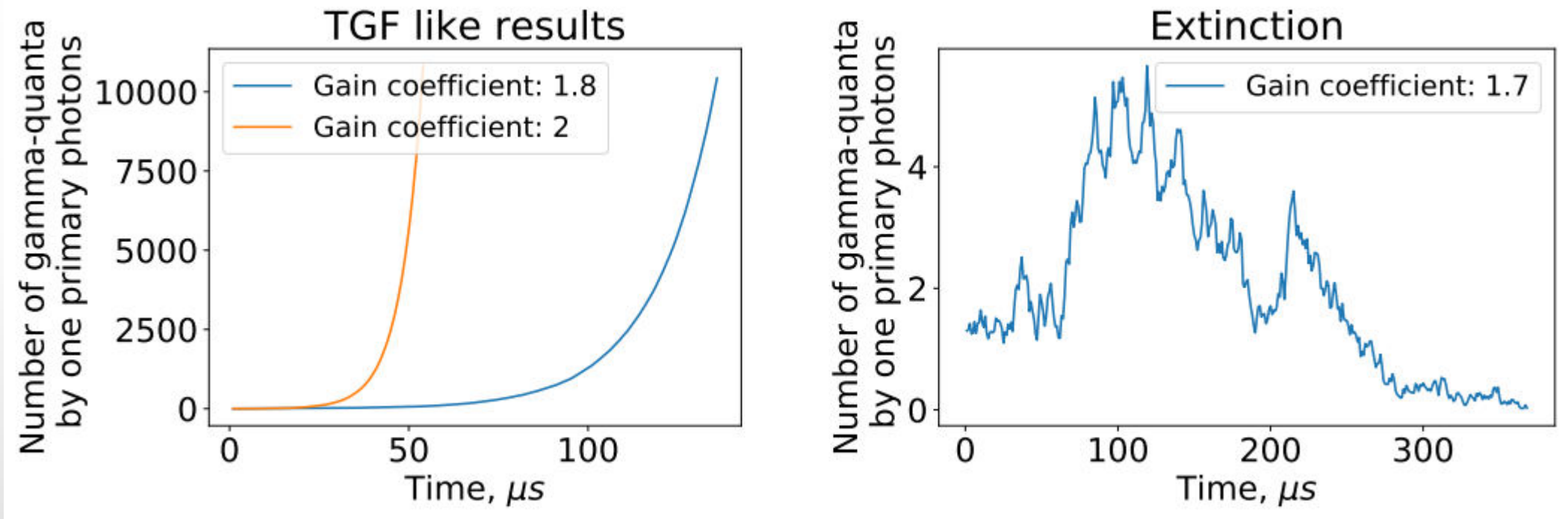
Help&Version

Start simulation

Window



Monte Carlo simulation results



Gain coefficient = local multiplication factor



Completely random reactor model assumptions

- The electric field is completely random, which makes gamma production isotropic.
- There is noncritical electric field between critical cells, consequently, there is only gamma exchange between cells.
- Critical electric field is the same everywhere in the cloud.
- All gammas have the same energy.
- Runaway electrons have the same energy.
- Air density is the same throughout the cloud.
- Gammas leave the system in two ways: fly out of the thundercloud or produce a runaway electron.



Equation

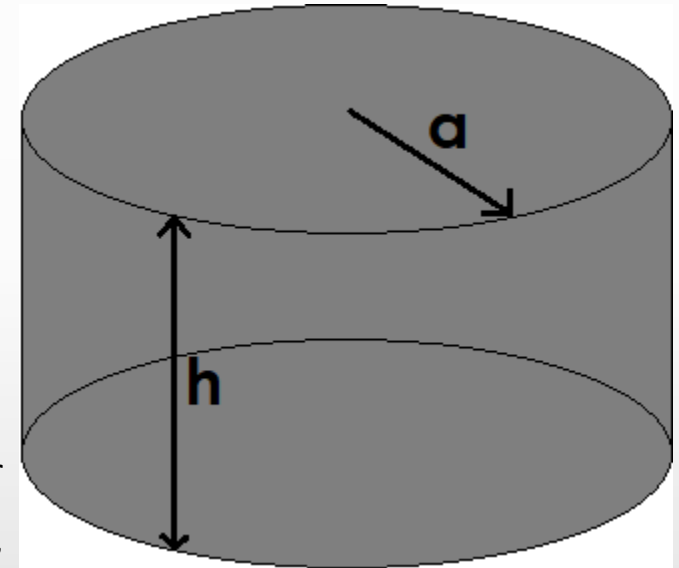
Under such assumptions gamma dynamics can be described with the following equation:

$$D\Delta n - c\Sigma n + \nu c\Sigma n = \frac{\partial n}{\partial t}$$

$$n|_{z=0,h} = 0$$

$$n|_{r=a} = 0$$

Here $n(\vec{r}, t)$ - gamma concentration, $D = \frac{c\lambda}{3}$ - diffusion coefficient, λ - mean free path length for gammas, $\Sigma = \frac{1}{\lambda_{\gamma \rightarrow e^-}}$ - mean macroscopic cross-section of runaway electron production by gamma, ν - local multiplication factor.





Local multiplication factor

Local multiplication factor is mean number of gammas generated by one gamma in a cell:

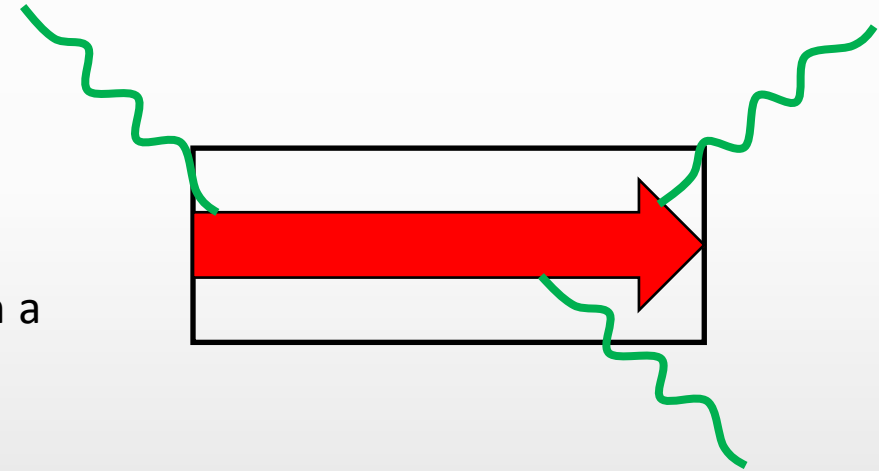
$$v = \text{RREA FORMATION PROBABILITY} \cdot \frac{\lambda_{\text{RREA}}}{\lambda_{e \rightarrow \gamma}} \cdot \left(\exp\left(\frac{L}{\lambda_{\text{RREA}}}\right) - 1 \right)$$

L – cell length.

RREA FORMATION PROBABILITY – average probability of RREA creation in a critical cell by gamma, also considering electric field geometry.

$\lambda_{\text{RREA}} = \frac{22mc^2}{eE}$ - Gurevich characteristic length of avalanche exponential growth (m, e – electron mass and charge, E – mean critical electric field).

$\lambda_{e \rightarrow \gamma}$ - gamma production by runaway electrons length.





The equation solution

The first member of the equation solution dominates:

$$n(r, z, t) = n_0 \cdot J_0 \left(\frac{2.405 \cdot r}{a} \right) \sin \left(\frac{\pi z}{h} \right) \cdot e^{\varepsilon t},$$

where $\varepsilon = \frac{\lambda c}{3} \left(\frac{3(\nu-1)}{\lambda \lambda_{\gamma \rightarrow e^-}} - \left(\frac{2.405}{a} \right)^2 - \left(\frac{\pi}{h} \right)^2 \right)$ - global multiplication factor.

Consequently, thunderstorm gamma flux comes out as follows:

$$|\Phi(r,t)| \Big|_{z=0,h} = \frac{\lambda c}{3} \frac{\partial n(r,z,t)}{\partial z} \Big|_{z=0,h} = \frac{\pi \lambda c}{h} n_0 \cdot J_0 \left(\frac{2.405 \cdot r}{a} \right) \cdot e^{\varepsilon t}$$



Reactor explosion criteria

Reactor becomes critical when number of gammas grows exponentially, or in other words, when $\varepsilon > 0$. Consequently, there is a following **reactor explosion criteria**:

$$\nu > \frac{\lambda\lambda_{\gamma \rightarrow e^-}}{3} \left(\left(\frac{2.405}{a} \right)^2 + \left(\frac{\pi}{h} \right)^2 \right) + 1$$

Or, considering the formula for local multiplication factor:

$$\text{RREA FORMATION PROBABILITY} \cdot \frac{22mc^2}{eE\lambda_{e^- \rightarrow \gamma}} \cdot \left(\exp\left(\frac{L}{\lambda_{\text{RREA}}}\right) - 1 \right) > \frac{\lambda\lambda_{\gamma \rightarrow e^-}}{3} \left(\left(\frac{2.405}{a} \right)^2 + \left(\frac{\pi}{h} \right)^2 \right) + 1$$

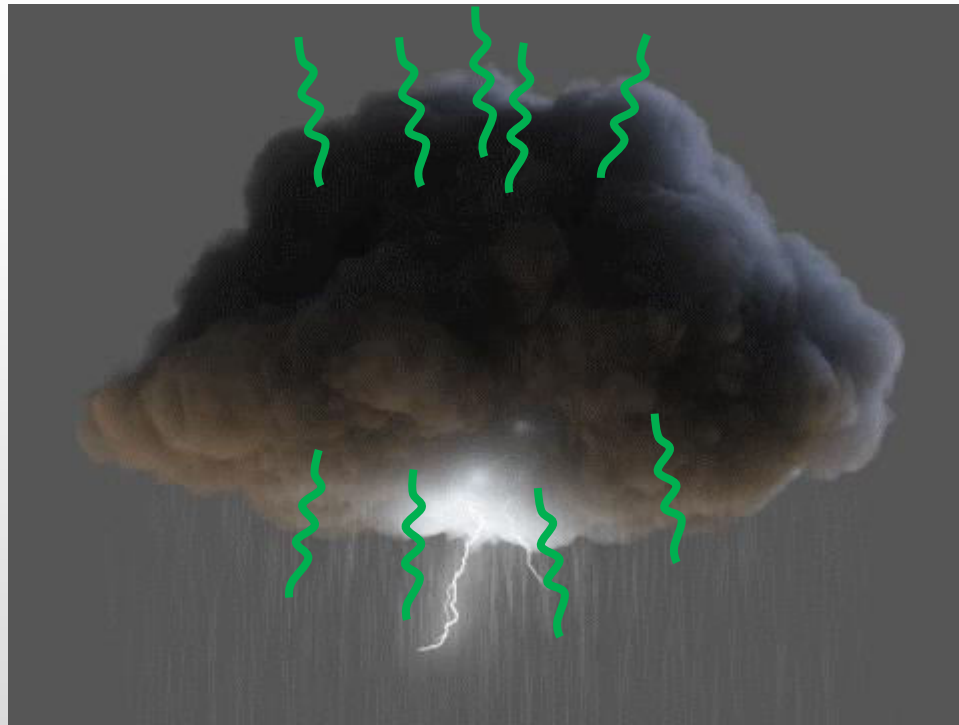




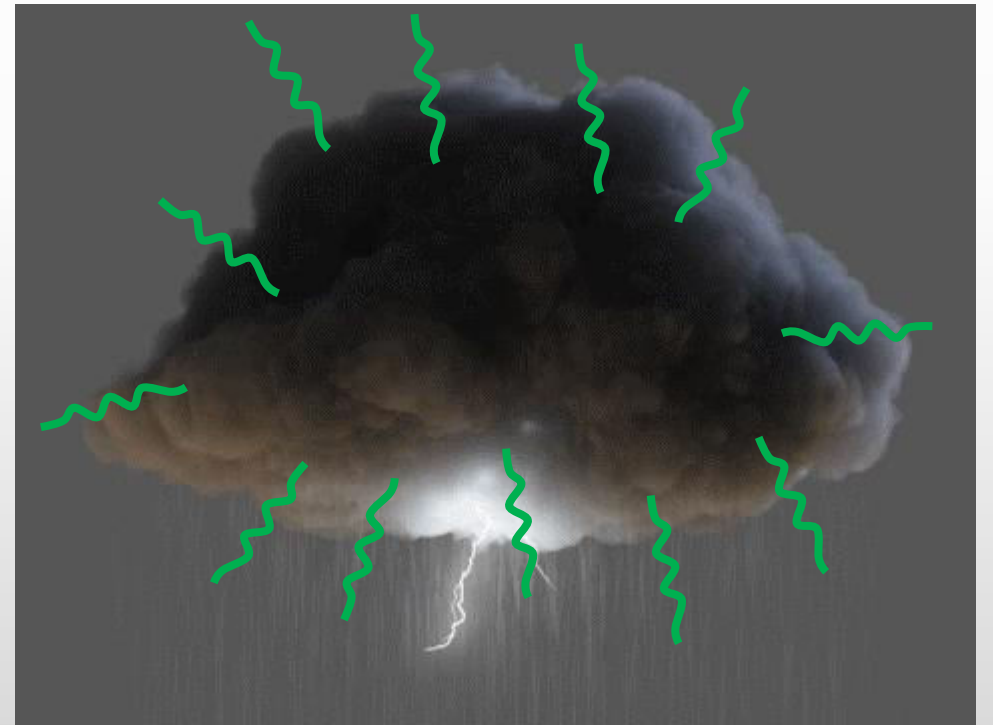
Reactor thundercloud gamma radiation

Reactor thundercloud gamma radiation is quasi-isotropic

Gurevich or Dwyer thundercloud



Reactor thundercloud





Gamma flux dependence on time

Considering cosmic ray flux gamma concentration dynamics is following:

$$dn = \frac{\partial n_{cosmic}}{\partial t} dt + n \cdot \varepsilon \cdot dt$$

$$n(0) = n_0$$

Here n – gamma concentration, ε – global multiplication factor.

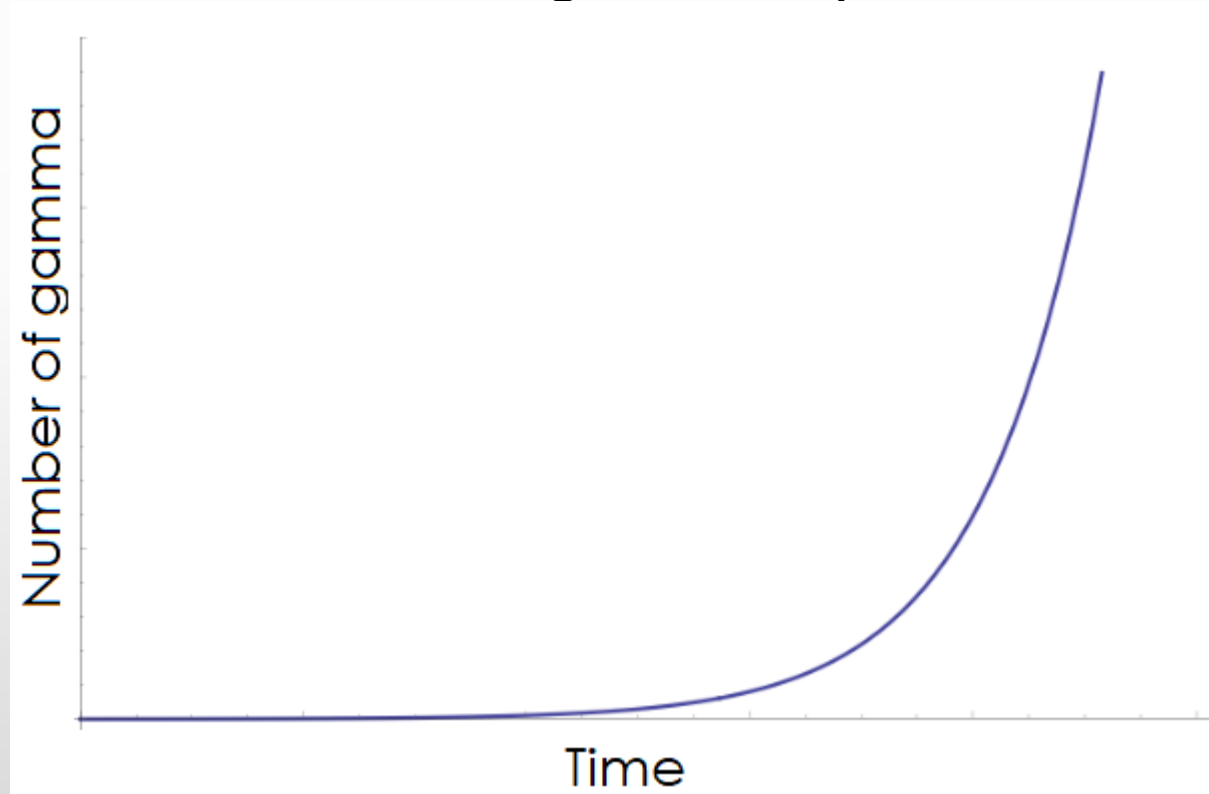
Consequently:

$$n(t) = n_0 e^{\varepsilon t} + \frac{\partial n_{cosmic}}{\partial t} \frac{e^{\varepsilon t} - 1}{\varepsilon}$$



Gamma flux dependence on time

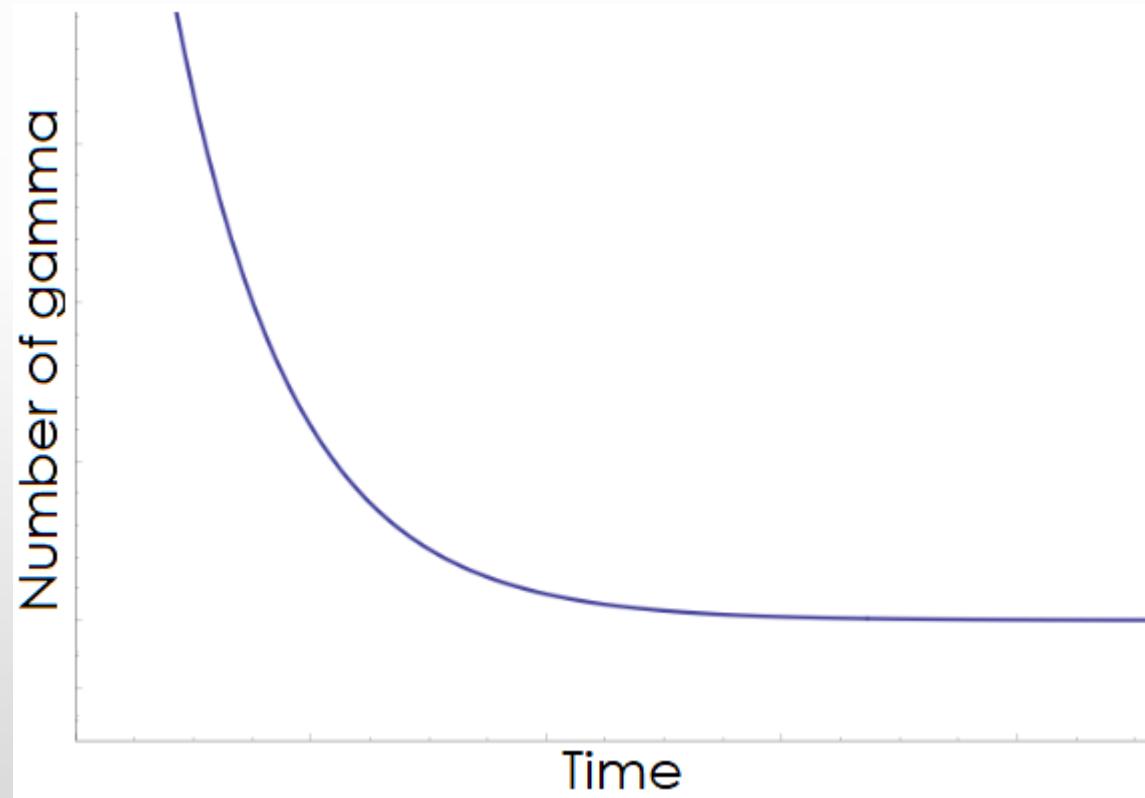
If $\varepsilon > 0$ then there is the following time dependence:





Gamma flux dependence on time

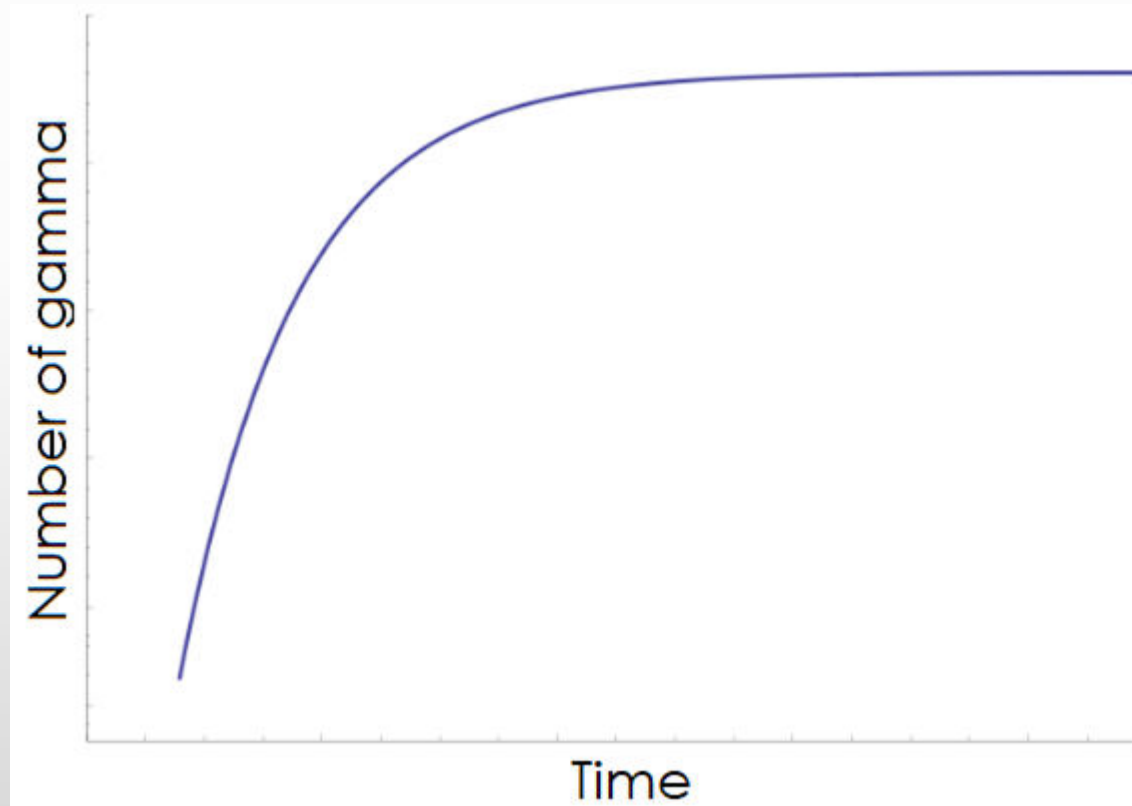
If $\varepsilon < 0$, $\frac{\partial n_{cosmic}}{\partial t} < \varepsilon n_0$ then there is the following time dependence:





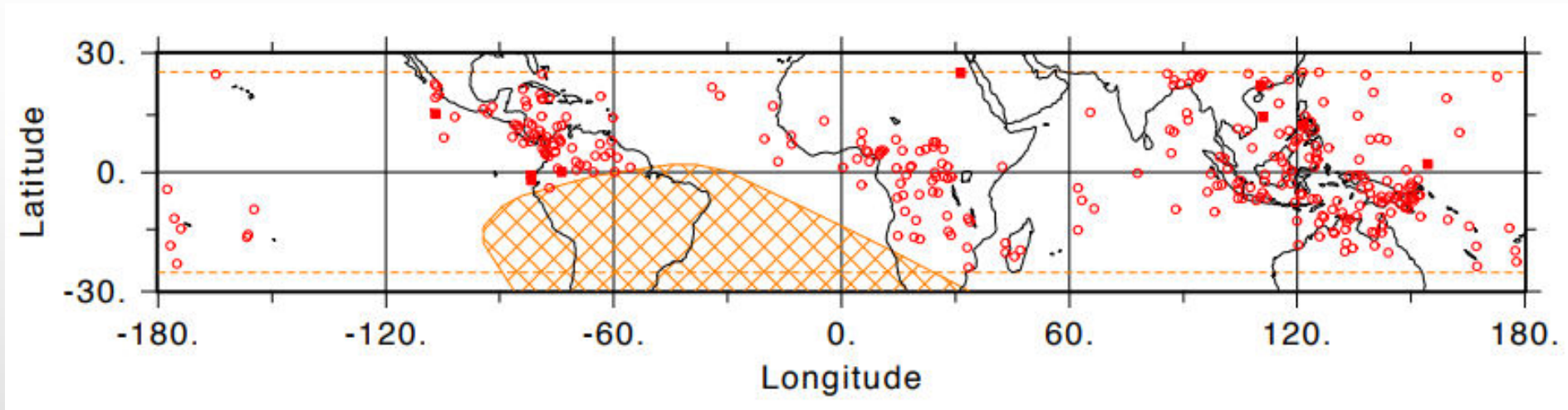
Gamma flux dependence on time

If $\varepsilon < 0$, $\frac{\partial n_{cosmic}}{\partial t} > \varepsilon n_0$ then there is the following time dependence:



TGF geography

In reactor model gamma dynamics depends on thundercloud size and electric field value. Consequently, TGF requires huge thunderclouds.

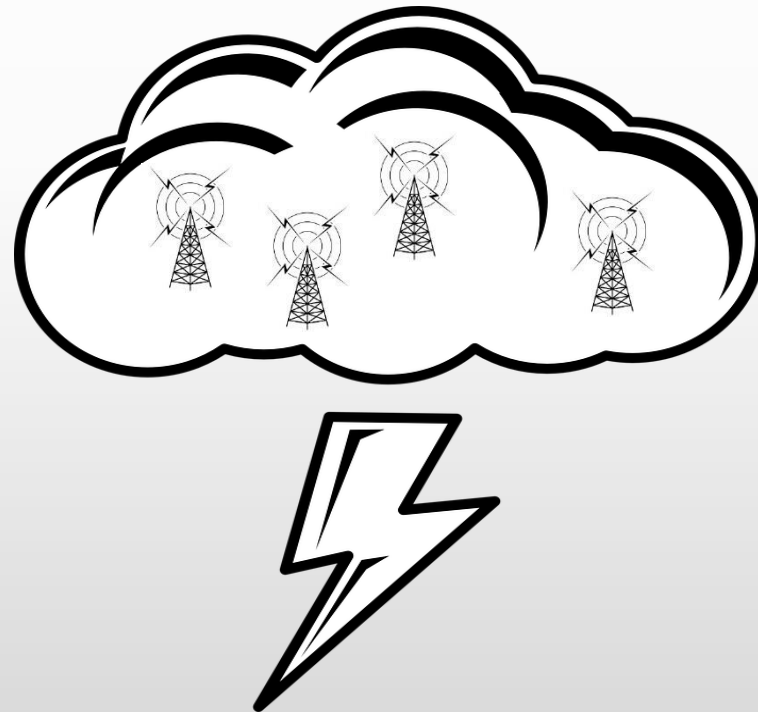


[Briggs - 2013 - Fermi GBM Observations of Terrestrial Gamma-ray Flashes \(TGFs \)](#)



Reactor radio signal

RREA radio signal generation is quasi-uniform throughout the thundercloud in reactor thundercloud.



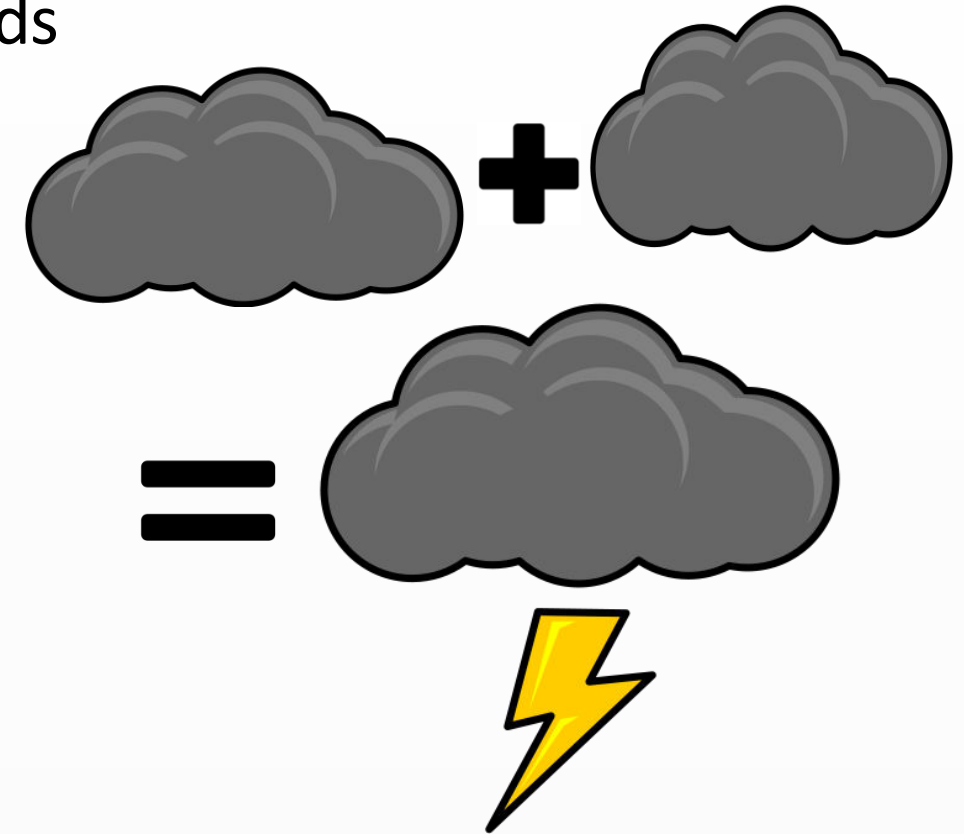


How to trigger TGF?

To burn only during hundreds of microseconds reactor explosion should be started initially with significant positive global multiplication factor.

List of TGF triggering hypothesizes:

- Undercritical thunderclouds collision
- Charge layers mixing by lightning
- Or simply low gamma fluxes are invisible from space.





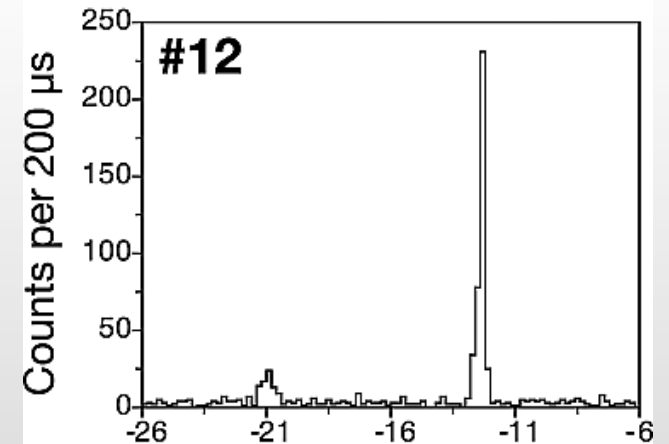
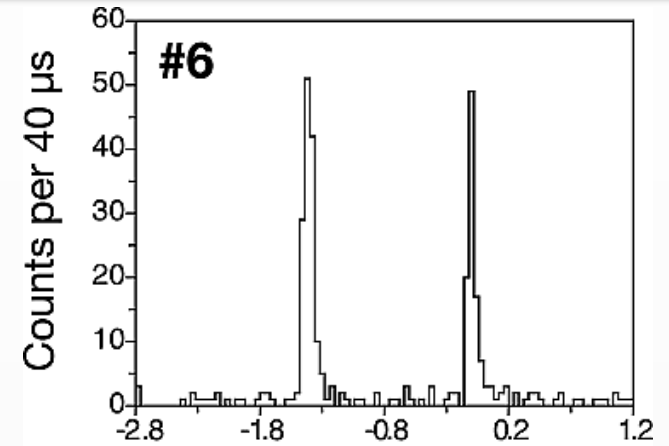
TGF dynamics in reactor model

TGF triggers discharges in cells.

Discharge at some cell can stop TGF making $\varepsilon < 0$.

That cell recharge can launch TGF again.

Consequently, multi-TGF is possible in reactor model.



<https://agupubs.onlinelibrary.wiley.com/doi/full/10.1029/2010JA016084>



Which came first: TGF or lightning?

In reactor model (hypothetically):

- TGF can cause lightning by triggering discharges.
- TGF can be terminated by lightning.
- Lightning can cause TGF by mixing cloud layers.





Bibliography

- First results on terrestrial gamma ray flashes from the Fermi Gamma-ray Burst Monitor / Briggs M. S., Fishman G. J., Connaughton V. et al. // Journal of Geophysical Research: Space Physics. — Vol. 115, no. A7. — URL: <https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2009JA015242>.
- Gurevich A.V., Milikh G.M., Roussel-Dupre R. Runaway electron mechanism of air breakdown and preconditioning during a thunderstorm // Physics Letters A. — 1992. — Vol. 165, no. 5. — Pp. 463 – 468. — URL: <http://www.sciencedirect.com/science/article/pii/037596019290348P>.
- R. Dwyer J. A fundamental limit on electric fields in air // Geophysical Research Letters. — Vol. 30, no. 20. — URL: <https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2003GL017781>.
- Gurevich Aleksandr V, Zybin Kirill P. Runaway breakdown and electric discharges in thunderstorms // Physics-Uspekhi. — 2001. — Vol. 44, no. 11. — P. 1119. — URL: <http://stacks.iop.org/1063-7869/44/i=11/a=R02>.
- Geant4—a simulation toolkit / S. Agostinelli, J. Allison, K. Amako et al. // Nuclear Instruments and Methods in Physics Research Section A: Accelerators

MIPT at a glance

Rankings

#48

THE
Physics

#67

THE Computer
Science

#42

QS Physics
& Astronomy

Alumni



Yuriy Baturin

Pilot astronaut, Hero of the Russian Federation



Alexander Kaleri

Pilot astronaut, Hero of the Russian Federation



Konstantin Novoselov

Nobel prizeman



Andrei Geim

Nobel prizeman



David Yan

Founder and Director of the board of ABBYY



Sergey Belousov

Founder and CEO of Acronis

Numbers

Founded in **1951**



Nobel prizemen among professors and alumni

80 Labs on campus

7132

Students

Phystech Schools



Radio Engineering and Computer Technology



Fundamental and Applied Physics



Aerospace Technology



Applied Mathematics and Informatics



Biological and Medical Physics



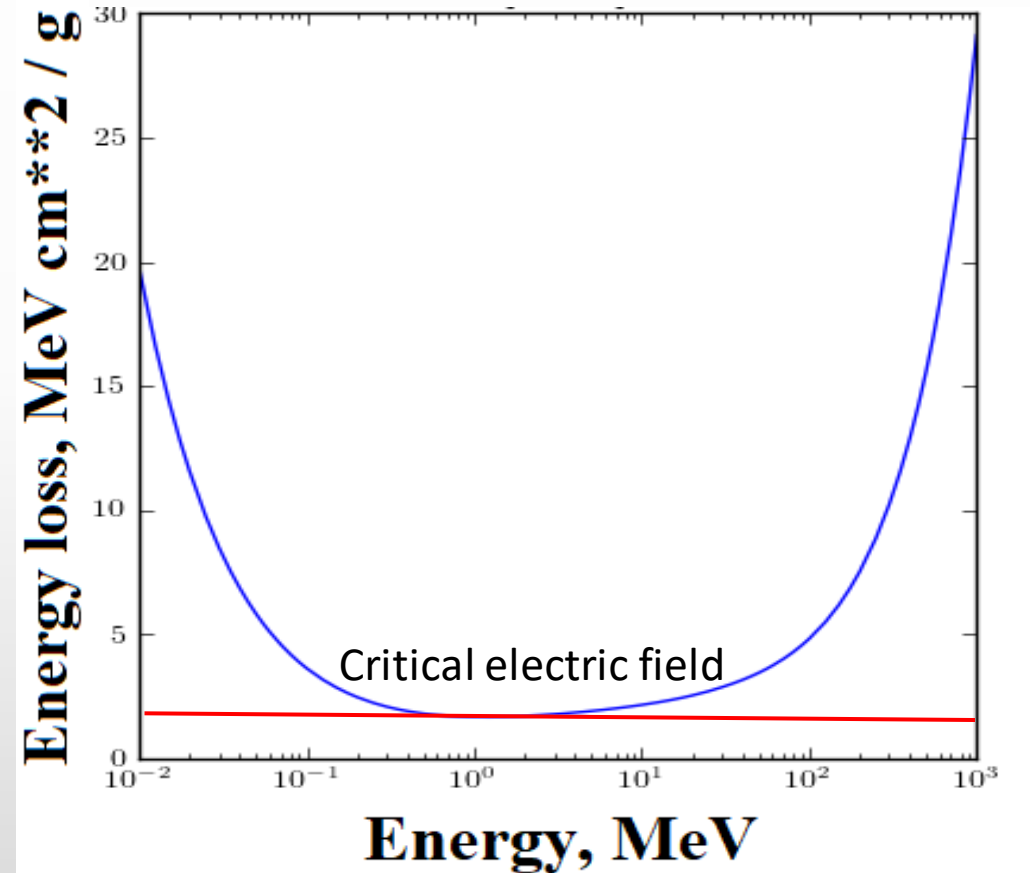
Electronics, Photonics and Molecular Physics



Runaway breakdown (Gurevich 1992)

Electric field might give to relativistic electron more energy than it wastes on interaction with air.

Such electric field is called **critical electric field**, accelerated electrons are called **runaway electrons**.





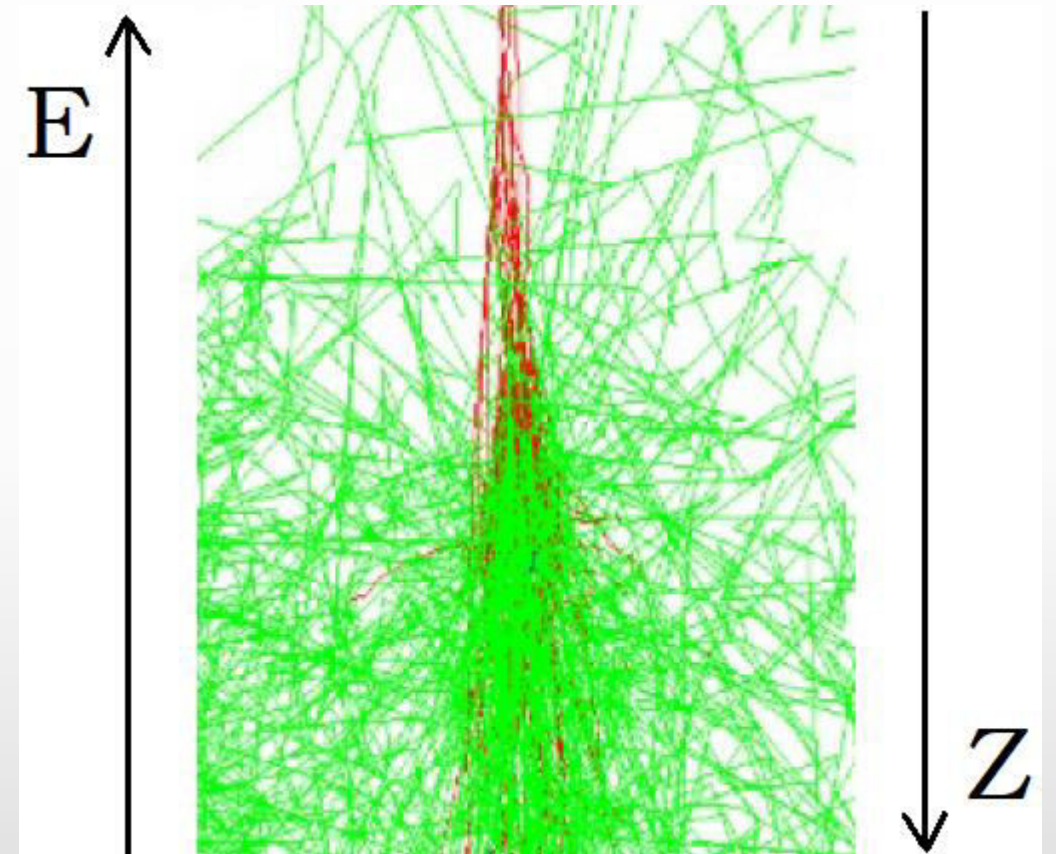
Relativistic runaway electron avalanche

Runaway electrons produce new runaway electrons by collision with air molecules' electrons.

The law of the **RREA** growth:

$$N(\mathbf{z}) = N_0 \cdot e^{\frac{\mathbf{z}}{l_a}}$$

$$l_a \sim 50 \text{ m}$$



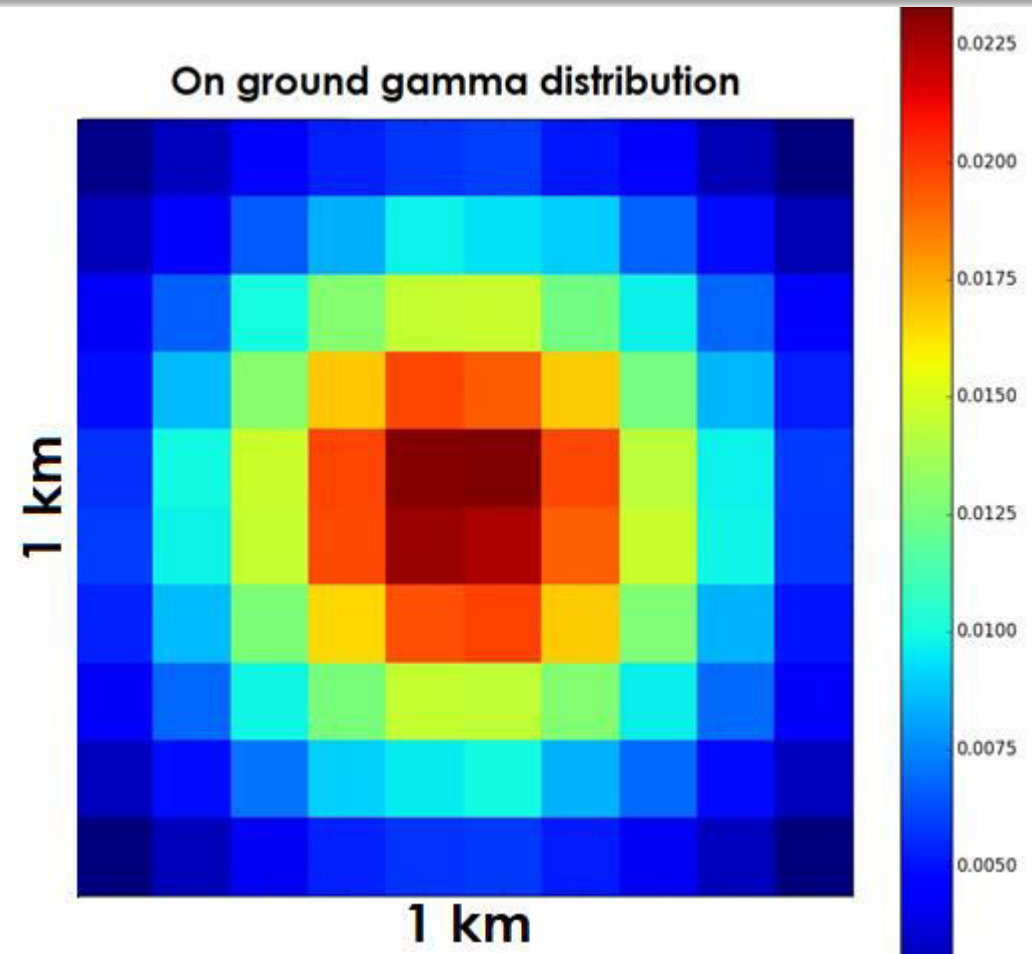
Geant4 simulation of RREA. **Green** – gammas, **red** – electrons.



Reactor thundercloud gamma radiation

On-ground quasi-isotropy hypothesis experimental observation is complicated due to air attenuation.

From 1 km isotropic gamma source looks like looks like directed source with Compton scattering.



Isotropic source Geant4 simulation



Reactor thundercloud gamma radiation

