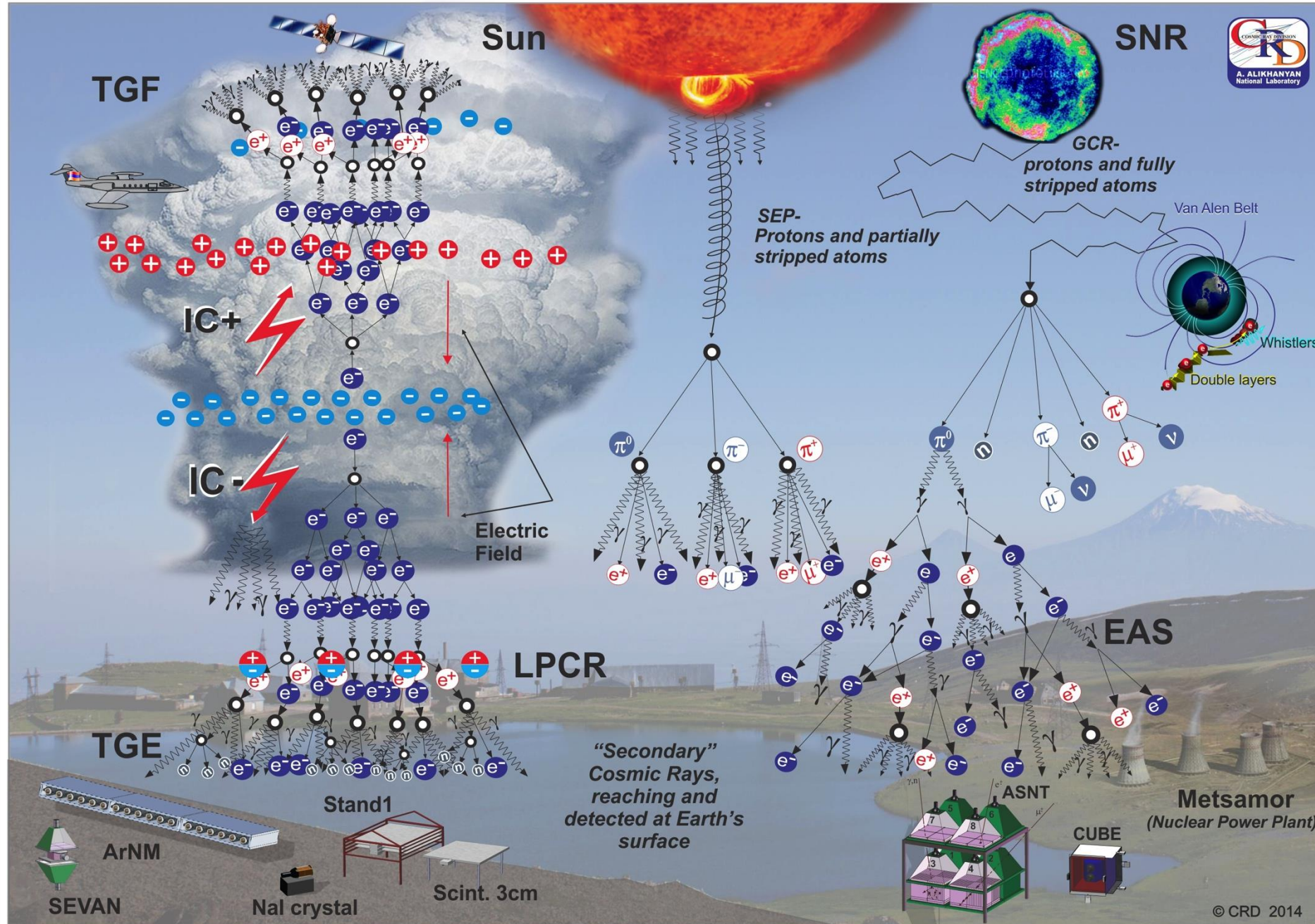


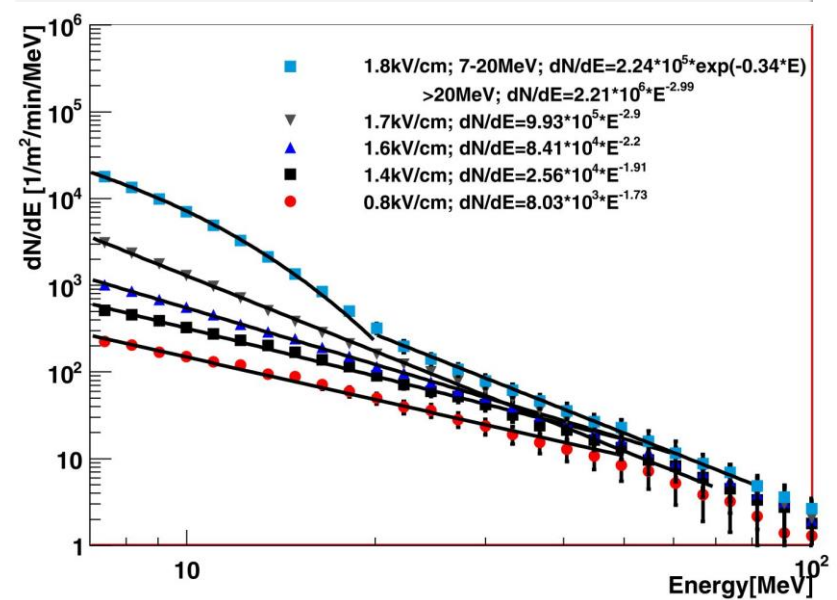
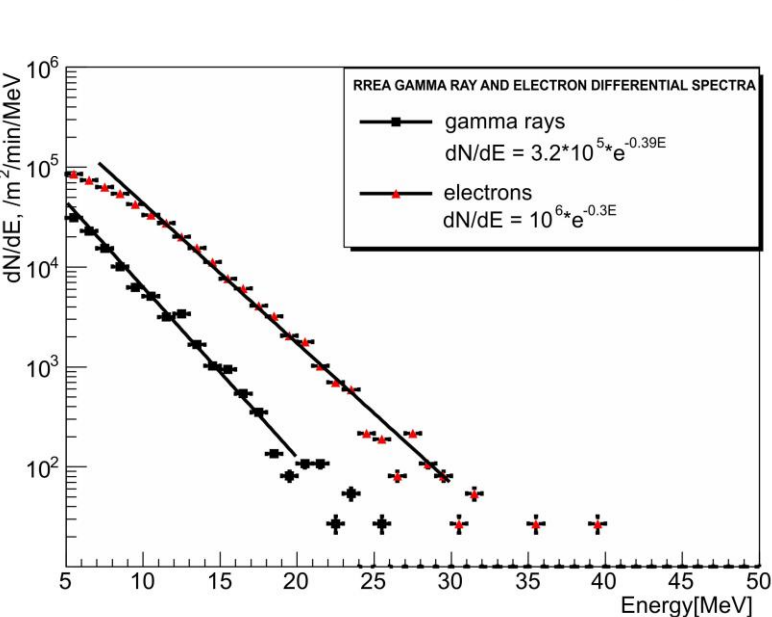
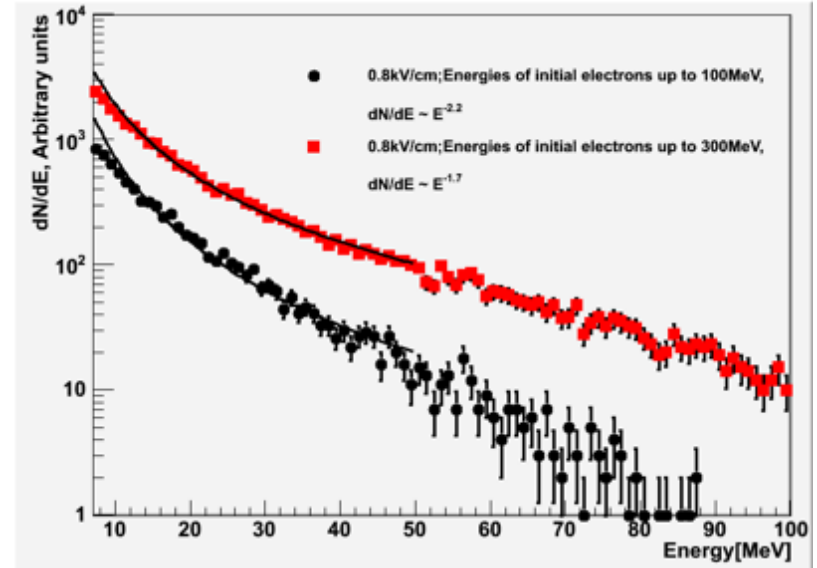
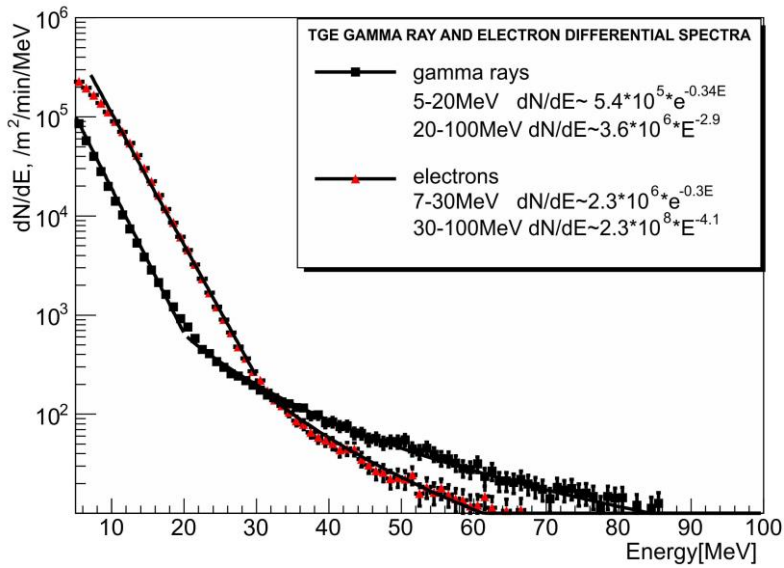
Origin of Secondary Cosmic Rays

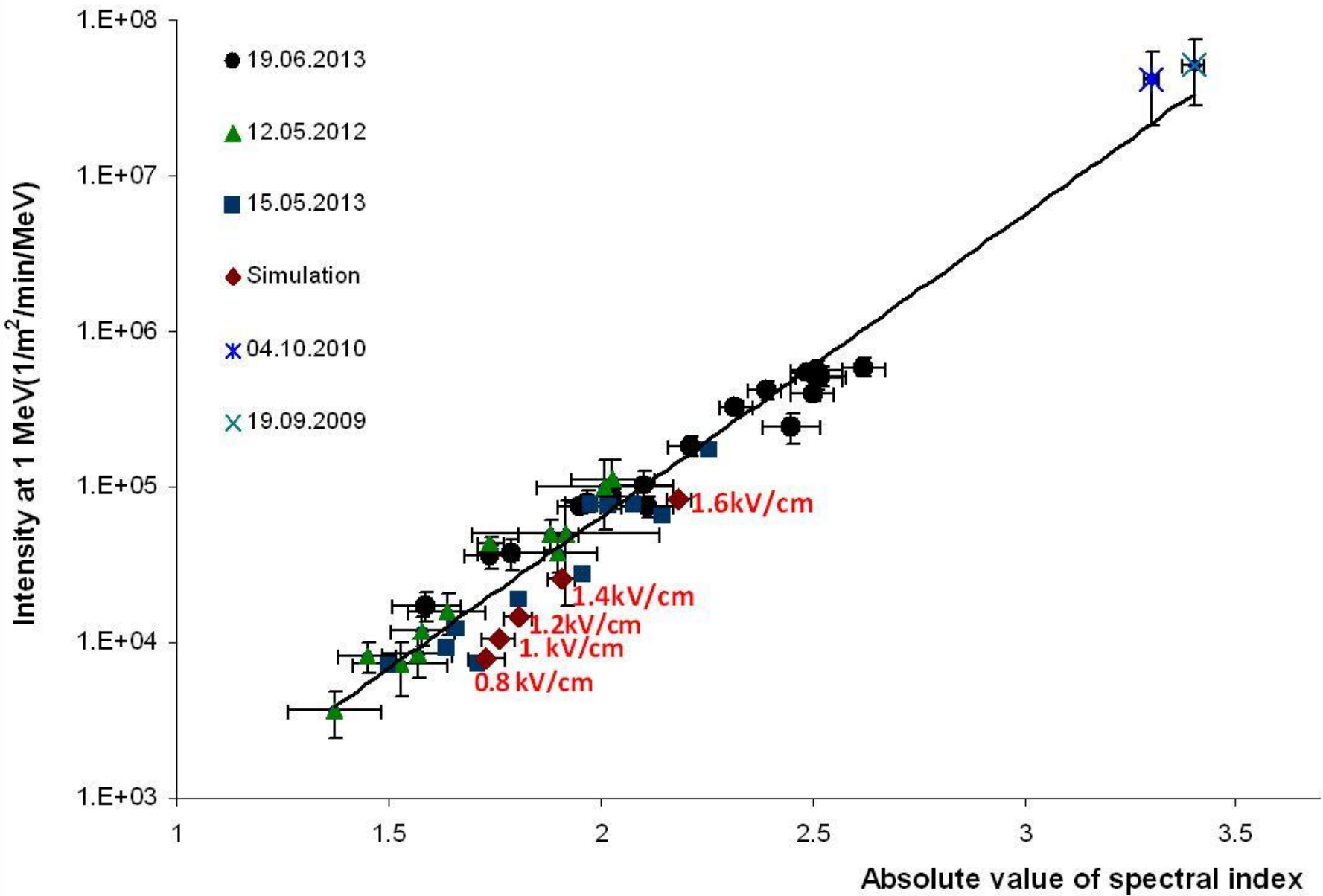


Discussions

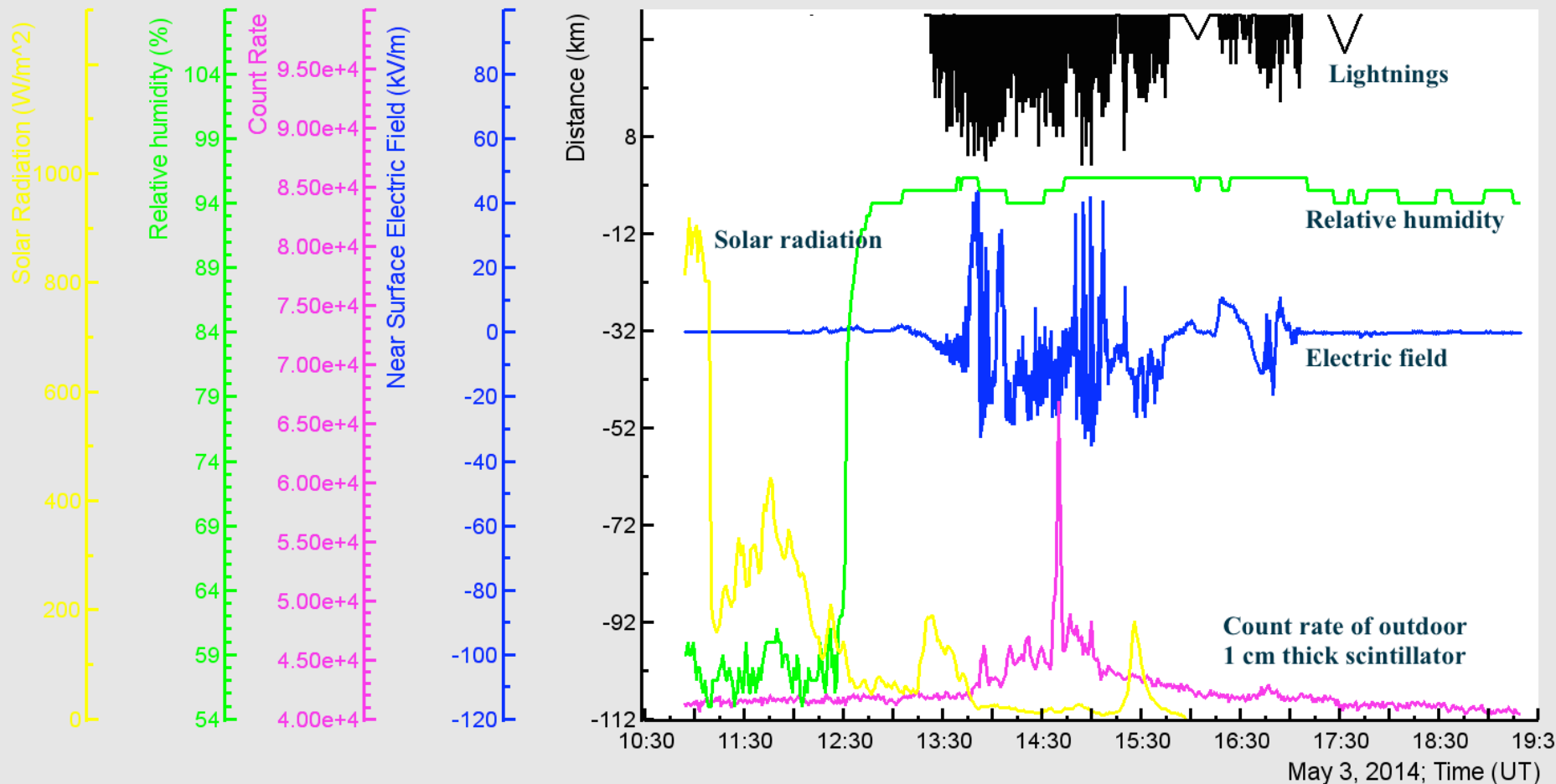
- TGEs and TGFs – symmetric processes?
- Lower dipole – upper dipole (positive – negative electric field); electrons accelerated downwards – upwards; IC- lightnings; IC+ lightnings;
- TGE – need developed LPCR – transient phenomenon; needs high RE, washed down by rain (origin of Long TGE) Observed and proved; TGF – dipole always present in thundercloud;
- RREA - observed in TGE, however – very rare process – manifestation – huge enhancements of flux, ECSses, exponential spectra; Don't observed in TGF; don't provide enough seed electrons if not include unproven exotic processes;
- MOS – power law energy spectra – observed in TGE; modest enhancement – most of TGE events; do not considered in TGF;
- Seeds: TGE – CR electrons – simulated agree with observations; TGF -the thunderstorm must have produced about 10^{17} high-energy electrons (average energy is 7 MeV), in order to account for the fluence of gamma rays recorded at the spacecraft many hundreds of kilometers away,
- Particle fluxes and lightnings – no direct causal relation; occurred after TGE and TGF.
- High energy physics in atmosphere - perspectives for coming years:
- Satellite missions;
- compare observed TGE characteristics with TGFs: will be it possible to recover individual, not cumulative TGF energy spectra? Compare spectral indices? Why TGF spectra are nor exponential if origin is avalanche process?

2-component model of TGE: GEANT4 simulation of TGE: Uniform Electrical fields of 0.8 – 1.8 kV/m, started from 5000 m till 3350 Observation on 3200. Seeds: CR ambient electrons

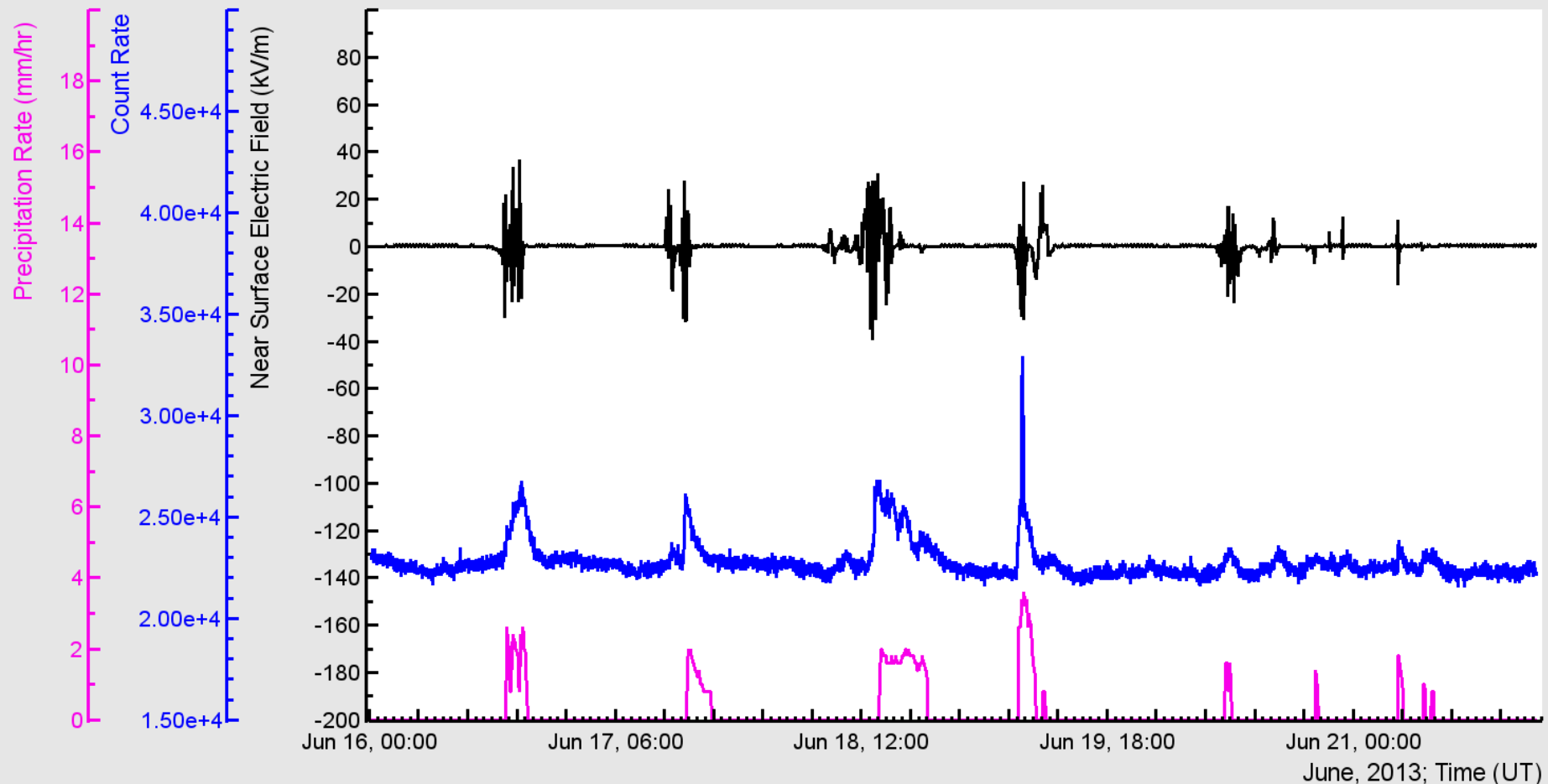




TGE is coincided with rise of Humidity – LPCR development

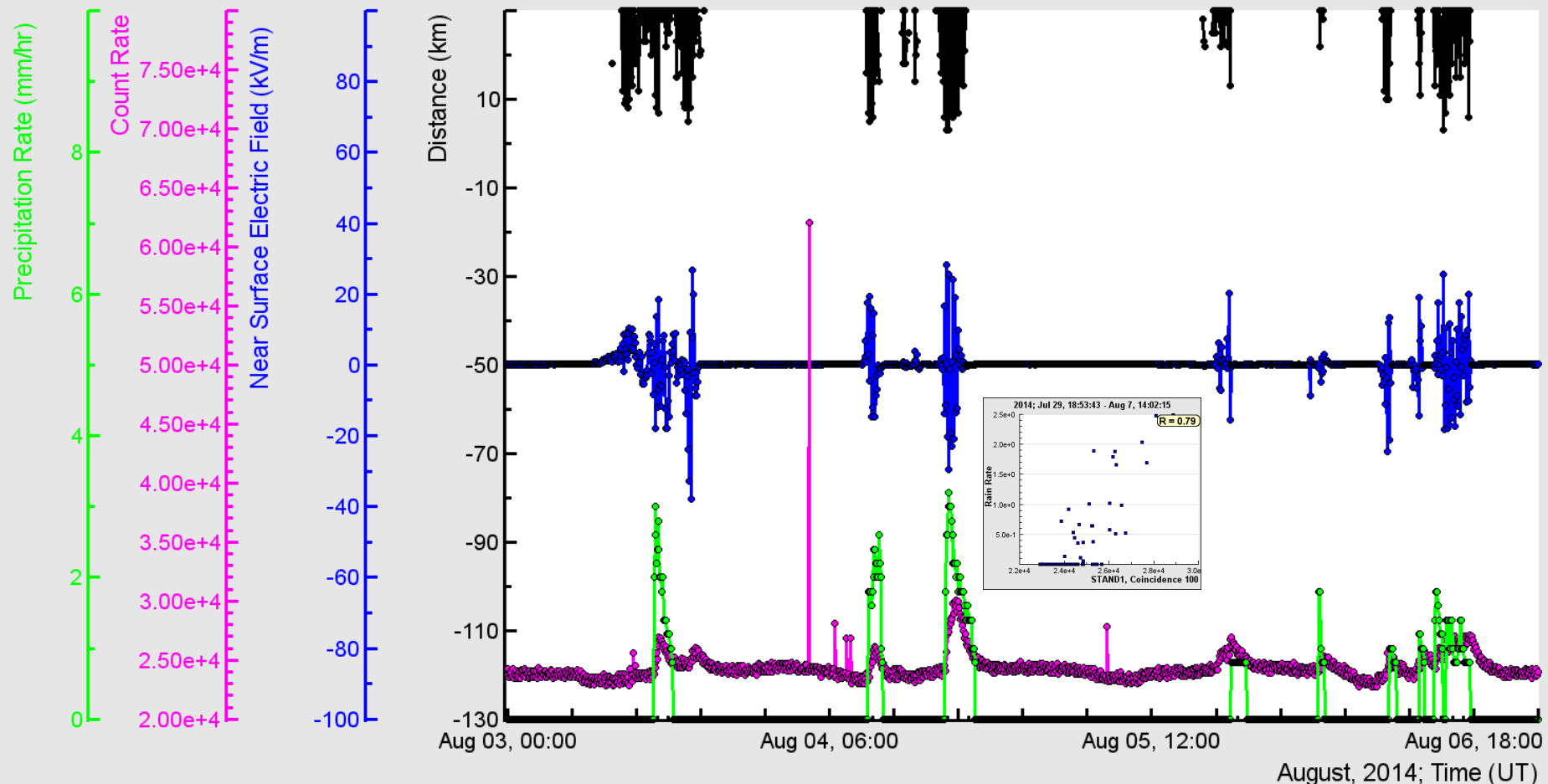


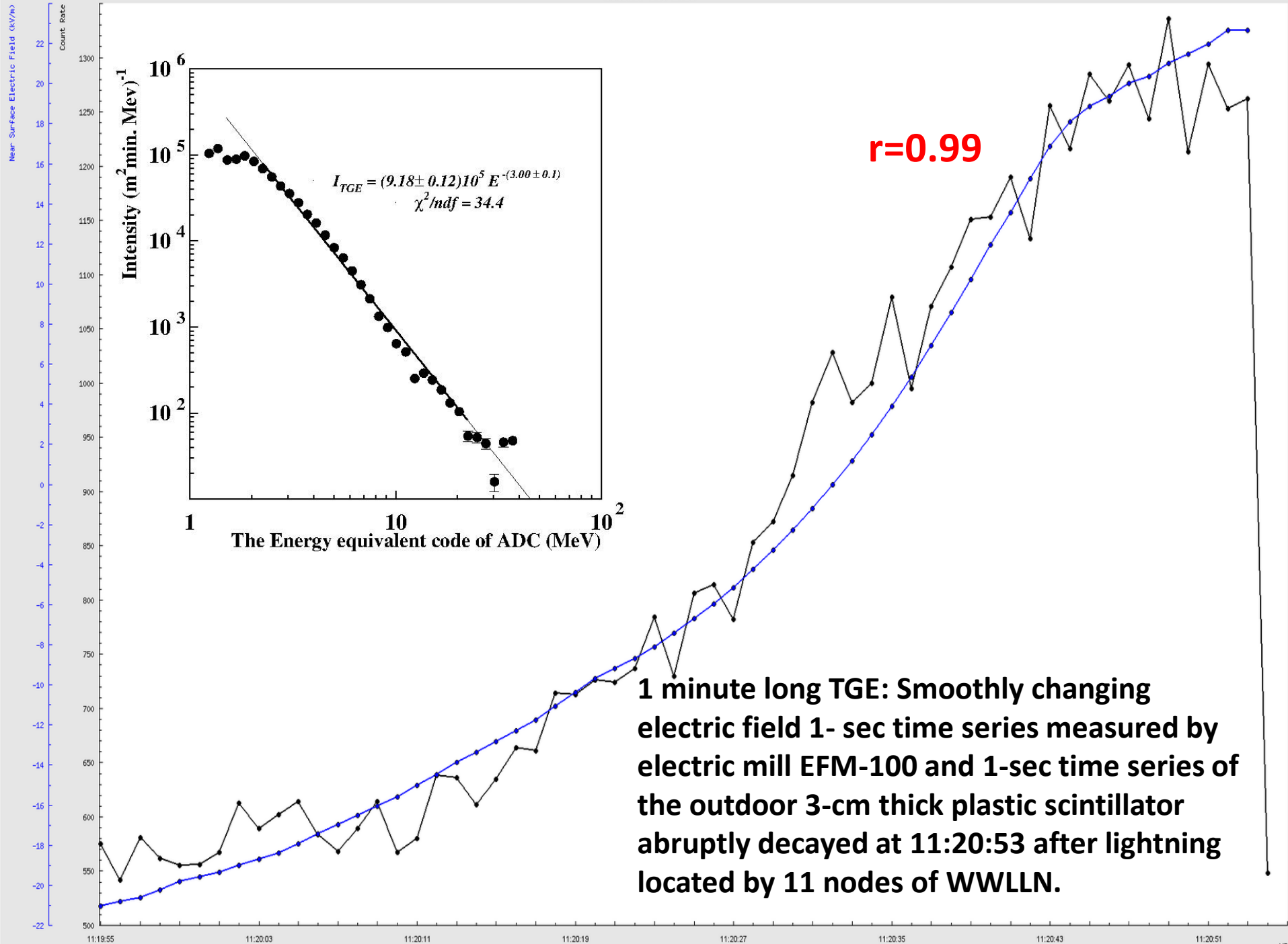
Charge reside on rain droplets making LPCR and- lower dipole accelerated electrons downward



After finishing of rain TGE fast decays

- rain wash down LPCR and lower dipole decays as well





**19 October 2013 TGE, 1-sec fluxes at 11:20:49 –
11:20:53 – compared with commulative flux AGILE
TGFs**

	Mean (CR) count rate 1/sec	TGE peak	TGE flux = PEAK- CR	Detector area (m²)	Det. eff.	TGE Flux** (1/cm²sec)
ScI	689+/- 25	1450	50,000	0.0135	0.8	7
3cm thick sc.	532+/- 23	1320	532	1	0.02	4
60 cm thick sc.***	3080+/- 55	10054	3080	1	0.20	1.7
NaI	75+/- 10	345	2340	0.032	0.8	1.2
AGILE 308 TGFs						17

Marisaldi, M., et al. (2014), Properties of terrestrial gamma ray flashes detected by AGILE MCAL below 30 MeV *J. Geophys. Res. Space Physics*, *119*, 1337–1355

2014 Aug 24 09:30:00 CUT

2014 Aug 23 10:30:00 CUT



Nearby CG; EFM ~0.1-0.2

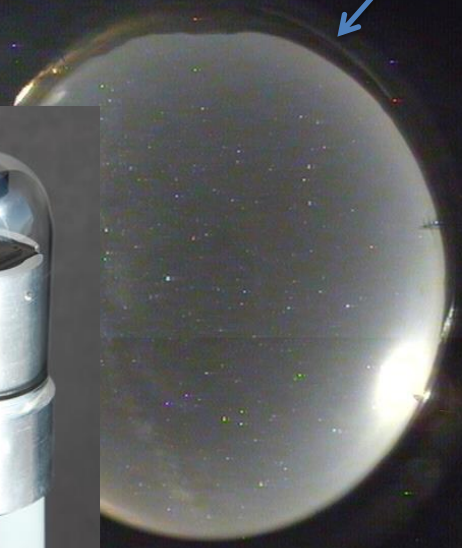
EFM < +/- 10 kV/m

EFM > 20 kV/m

2014 Aug 2 18:00:00 CUT

CUT

2014 Jun 2 21:00:00 CUT



TEPA-2015?

- Comments on TEPA-2014;
 - Students support;
 - Proceedings? End of 2014.
-
- TEPA-2015:
 - Nor – Amberd? Moscow?
 - Dates? August?